

Package ‘CRAC’

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Type Package

Title Cosmology R Analysis Code

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Description R functions for cosmological research.

The main functions are similar to the python library, cosmology.

License GPL (>= 2)

References [H99] Hogg 1999, <http://arxiv.org/abs/astro-ph/9905116>

[PDG] Particle Data Group,
http://pdg.lbl.gov/2013/reviews/contents_sports.html

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distance.angular *compute the angular diameter distance [Mpc/h]*

Description

ONLY FOR flat universe, $\Omega_k = 0$.

Usage

```
distance.angular(z, cosmo, z0 = 0, ...)
```

Arguments

<code>z</code>	redshift upper limit
<code>cosmo</code>	cosmological parameter list
<code>z0</code>	redshift lower limit
<code>...</code>	pass to integrate() to control integration accuracy.

Value

Angular diameter distance from $z_0 (= 0)$ to z [Mpc/h]

References

Equation (18) in [H99]

See Also

[distance.comoving](#)

Examples

```
distance.angular(0.1,parameter.fidcosmo)
distance.angular(0.3,list(omegaM0=0.272,omegaL0=0.728,omegaK=0.0,h=0.704))
```

distance.comoving	<i>compute the comoving distance (line-of-sight) [Mpc/h]</i>
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Description

compute the comoving distance (line-of-sight) [Mpc/h]

Usage

```
distance.comoving(z, cosmo, z0 = 0, ...)
```

Arguments

z	redshift upper limit
cosmo	cosmological parameter list
z0	redshift lower limit
...	pass to integrate() to control integration accuracy.

Value

Comoving distance from $z_0 (= 0)$ to z to z [Mpc/h]

References

Equation (15) in [H99]

Examples

```
distance.comoving(0.2,parameter.fidcosmo,z0=0.3)
sapply(seq(0,1,0.1),function (x) distance.comoving(x,parameter.fidcosmo))
```

distance.luminosity	<i>Compute the luminosity distance [Mpc/h]</i>
---------------------	--

Description

Compute the luminosity distance [Mpc/h]

Usage

```
distance.luminosity(z, cosmo, z0 = 0, ...)
```

Arguments

<code>z</code>	redshift upper limit
<code>cosmo</code>	cosmological parameter list
<code>z0</code>	redshift lower limit
<code>...</code>	pass to <code>integrate()</code> to control integration accuracy.

Value

Luminosity distance from $z_0 (= 0)$ to z [Mpc/h]

References

Equation (20) in [H99]

See Also

[distance.angular](#), [distance.comoving](#)

Examples

```
distance.luminosity(0.1,parameter.fidcosmo)
```

`distance.transverse` *compute the comoving distance (transverse) [Mpc/h]*

Description

compute the comoving distance (transverse) [Mpc/h]

Usage

```
distance.transverse(z, cosmo, z0 = 0, ...)
```

Arguments

<code>z</code>	redshift upper limit
<code>cosmo</code>	cosmological parameter list
<code>z0</code>	redshift lower limit
<code>...</code>	pass to <code>integrate()</code> to control integration accuracy.

Value

Comoving distance from $z_0 (= 0)$ to z [Mpc/h]

References

Equation (16) in [H99]

See Also[distance.comoving](#)**Examples**

```
distance.transverse(0.1,parameter.fidcosmo)
```

eZ*compute $E(z)$ of given cosmology*

Description

Compute the $E(z) = H(z)/H_0$, or

$$E(z) \equiv \sqrt{\Omega_M(1+z)^3 + \Omega_k(1+z)^2 + \Omega_\Lambda(1+z)^{1+w}}, \text{where we omit the radiation component.}$$

Usage

```
eZ(z, cosmo)
```

Arguments

<code>z</code>	Redshift
<code>cosmo</code>	cosmology parameter list, contains 'omegaM0', 'omegaL0', 'omegaK'

Value

The dimensionless Hubble constant $H(z)/H_0$

References

Equation (14) in [H99]

See Also[eZ2,parameter.fidcosmo](#)**Examples**

```
eZ(1.2,parameter.fidcosmo)
```

eZ2*compute $E^2(z)$ of given cosmology***Description**

Compute the $E^2(z) = (H(z)/H_0)^2$

Usage

```
eZ2(z, cosmo)
```

Arguments

<code>z</code>	Redshift
<code>cosmo</code>	cosmology parameter list, contains 'omegaM0', 'omegaL0', 'omegaK'

Value

The dimensionless Hubble constant $H(z)/H_0$

References

Equation (14) in [H99]

See Also

[eZ](#)

Examples

```
eZ2(1.2,parameter.fidcosmo)
```

parameter.constant*Physical constants in SI units, source: [PDG]***Description**

Useful physical constants in SI units.

Usage

```
parameter.constant
```

Format

```
List of 8
$ c      : num 3e+08
$ me     : num 9.11e-31
$ h      : num 6.63e-34
$ mp     : num 1.67e-27
$ e      : num 1.6e-19
$ G      : num 6.67e-11
$ k      : num 1.38e-23
$ sigma.T: num 6.65e-29
```

parameter.DH

*Hubble Distance $D_H = c/H_0$ [Mpc/h]***Description**Hubble Distance $D_H = c/H_0$ [Mpc/h]**Usage**

parameter.DH

Format

num 2998

parameter.fidcosmo

*Fiducial cosmology parameter list***Description**

the parameter list is a crucial input for other calculations

Usage

parameter.fidcosmo

Format

```
List of 4
$ omegaM0: num 0.3
$ omegaL0: num 0.7
$ omegaK : num 0
$ h      : num 0.7
```

`parameter.unit` *useful conversion of astrophysical units*

Description

useful conversion of astrophysical units

Usage

`parameter.unit`

Format

```
List of 3
$ Mpc : num 3.09e+22
$ Msun: num 1.99e+30
$ eV   : num 1.6e-19
```

`rhoc` *calculate the critical density at redshift z*

Description

calculate the critical density at redshift z

Usage

`rhoc(z, cosmo)`

Arguments

<code>z</code>	redshift
<code>cosmo</code>	cosmological parameter list return the critical densith at given redshift in $h_0^2 \text{kg/m}^3]$

Examples

```
# get the critial density at z=0
rhoc(0,parameter.fidcosmo)
```

valid.cosmo	<i>validate the cosmological parameters are enough</i>
-------------	--

Description

validate the cosmological parameter list is complete for other usage. If missing data detected, it will be filled with fiducial value. Also the warning message is printed in red for *nix enviroment.

Usage

```
valid.cosmo(cosmo)
```

Arguments

cosmo	cosmological parameter list
-------	-----------------------------

Value

the cosmo list will be updated if important variables are missing

See Also

[parameter.fidcosmo](#)

Examples

```
# there are two typos in the cosmology parameter list
a <- list(omegaM=0.272,omegaL0=0.728,omegaK=0.0,h0=0.704)
valid.cosmo(a)
print(a)
```

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